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# nvector Documentation

*Release 0.6.0*

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[pypi package](#)  This is the documentation of **nvector** 0.6 for Python.

Bleeding edge available at: <https://github.com/pbrod/nvector>.

Official releases are available at: <http://pypi.python.org/pypi/nvector>.

Official homepage are available at: <http://www.navlab.net/nvector/>



# CHAPTER 1

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## Contents

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### 1.1 Introduction to Nvector

[pypi package 2.2.2](#)

Nvector is a suite of tools written in Python to solve geographical position calculations like:

- Calculate the surface distance between two geographical positions.
- Convert positions given in one reference frame into another reference frame.
- Find the destination point given start point, azimuth/bearing and distance.
- Find the mean position (center/midpoint) of several geographical positions.
- Find the intersection between two paths.
- Find the cross track distance between a path and a position.

### 1.2 Description

In this library, we represent position with an “n-vector”, which is the normal vector to the Earth model (the same reference ellipsoid that is used for latitude and longitude). When using n-vector, all Earth-positions are treated equally, and there is no need to worry about singularities or discontinuities. An additional benefit with using n-vector is that many position calculations can be solved with simple vector algebra (e.g. dot product and cross product).

Converting between n-vector and latitude/longitude is unambiguous and easy using the provided functions.

`n_E` is n-vector in the program code, while in documents we use `nE`. `E` denotes an Earth-fixed coordinate frame, and it indicates that the three components of n-vector are along the three axes of `E`. More details about the notation and reference frames can be found [here](#):

### 1.3 Documentation and code

Official documentation:

<http://www.navlab.net/nvector/>

<http://nvector.readthedocs.io/en/latest/>

**Kenneth Gade (2010):** A Nonsingular Horizontal Position Representation, The Journal of Navigation, Volume 63, Issue 03, pp 395-417, July 2010.

Bleeding edge: <https://github.com/pbrod/nvector>.

Official releases available at: <http://pypi.python.org/pypi/nvector>.

## 1.4 Installation

If you have pip installed and are online, then simply type:

```
$ pip install nvector
```

to get the lastest stable version. Using pip also has the advantage that all requirements are automatically installed.

You can download nvector and all dependencies to a folder “pkg”, by the following:

```
$ pip install --download=pkg nvector
```

To install the downloaded nvector, just type:

```
$ pip install --no-index --find-links=pkg nvector
```

## 1.5 Unit tests

To test if the toolbox is working paste the following in an interactive python session:

```
import nvector as nv
nv.test('--doctest-modules')
```

or

```
$ py.test --pyargs nvector --doctest-modules
```

at the command prompt.

## 1.6 Acknowledgement

The nvector package for Python was written by Per A. Brodkorb at FFI (The Norwegian Defence Research Establishment) based on the nvector toolbox for Matlab written by the navigation group at FFI.

Most of the content is based on the following article:

**Kenneth Gade (2010):** A Nonsingular Horizontal Position Representation, The Journal of Navigation, Volume 63, Issue 03, pp 395-417, July 2010.

Thus this article should be cited in publications using this page or the downloaded program code.

## 1.7 Getting Started

Below the object-oriented solution to some common geodesic problems are given. In the first example the functional solution is also given. The functional solutions to the remaining problems can be found [here](#).

### 1.7.1 Example 1: “A and B to delta”



Given two positions, A and B as latitudes, longitudes and depths relative to Earth, E.

Find the exact vector between the two positions, given in meters north, east, and down, and find the direction (azimuth) to B, relative to north. Assume WGS-84 ellipsoid. The given depths are from the ellipsoid surface. Use position A to define north, east, and down directions. (Due to the curvature of Earth and different directions to the North Pole, the north, east, and down directions will change (relative to Earth) for different places. A must be outside the poles for the north and east directions to be defined.)

**Solution:**

```
>>> import numpy as np
>>> import nvector as nv
>>> wgs84 = nv.FrameE(name='WGS84')
>>> pointA = wgs84.GeoPoint(latitude=1, longitude=2, z=3, degrees=True)
>>> pointB = wgs84.GeoPoint(latitude=4, longitude=5, z=6, degrees=True)
```

**Step1: Find p\_AB\_N (delta decomposed in N).**

```
>>> p_AB_N = pointA.delta_to(pointB)
>>> x, y, z = p_AB_N.pvector.ravel()
>>> valtxt = '{0:8.2f}, {1:8.2f}, {2:8.2f}'.format(x, y, z)
>>> 'Ex1: delta north, east, down = {}'.format(valtxt)
'Ex1: delta north, east, down = 331730.23, 332997.87, 17404.27'
```

**Step2: Also find the direction (azimuth) to B, relative to north:**

```
>>> azimuth = p_AB_N.azimuth_deg[0]
>>> 'azimuth = {0:4.2f} deg'.format(azimuth)
'azimuth = 45.11 deg'
```

**Functional Solution:**

```
>>> import numpy as np
>>> import nvector as nv
>>> from nvector import rad, deg
```

```
>>> lat_EA, lon_EA, z_EA = rad(1), rad(2), 3
>>> lat_EB, lon_EB, z_EB = rad(4), rad(5), 6
```

**Step1: Convert to n-vectors:**

```
>>> n_EA_E = nv.lat_lon2n_E(lat_EA, lon_EA)
>>> n_EB_E = nv.lat_lon2n_E(lat_EB, lon_EB)
```

**Step2: Find p\_AB\_E (delta decomposed in E).WGS-84 ellipsoid is default:**

```
>>> p_AB_E = nv.n_EA_E_and_n_EB_E2p_AB_E(n_EA_E, n_EB_E, z_EA, z_EB)
```

**Step3: Find R\_EN for position A:**

```
>>> R_EN = nv.n_E2R_EN(n_EA_E)
```

**Step4: Find p\_AB\_N (delta decomposed in N).**

```
>>> p_AB_N = np.dot(R_EN.T, p_AB_E).ravel()
>>> valtxt = '{0:8.2f}, {1:8.2f}, {2:8.2f}'.format(*p_AB_N)
>>> 'Ex1: delta north, east, down = {}'.format(valtxt)
'Ex1: delta north, east, down = 331730.23, 332997.87, 17404.27'
```

**Step5: Also find the direction (azimuth) to B, relative to north:**

```
>>> azimuth = np.arctan2(p_AB_N[1], p_AB_N[0])
>>> 'azimuth = {} deg'.format(deg(azimuth))
'azimuth = 45.11 deg'
```

See also Example 1 at [www.navlab.net](http://www.navlab.net)

### 1.7.2 Example 2: “B and delta to C”



A radar or sonar attached to a vehicle B (Body coordinate frame) measures the distance and direction to an object C. We assume that the distance and two angles (typically bearing and elevation relative to B) are already combined to the vector p\_BC\_B (i.e. the vector from B to C, decomposed in B). The position of B is given as n\_EB\_E and z\_EB, and the orientation (attitude) of B is given as R\_NB (this rotation matrix can be found from roll/pitch/yaw by using zyx2R).

Find the exact position of object C as n-vector and depth ( n\_EC\_E and z\_EC ), assuming Earth ellipsoid with semi-major axis a and flattening f. For WGS-72, use a = 6 378 135 m and f = 1/298.26.

**Solution:**

```
>>> import nvector as nv
>>> import numpy as np
>>> wgs72 = nv.FrameE(name='WGS72')
>>> wgs72 = nv.FrameE(a=6378135, f=1.0/298.26)
```

**Step 1: Position and orientation of B is given 400m above E:**

```
>>> n_EB_E = wgs72.Nvector(nv.unit([[1], [2], [3]]), z=-400)
>>> frame_B = nv.FrameB(n_EB_E, yaw=10, pitch=20, roll=30, degrees=True)
```

**Step 2: Delta BC decomposed in B**

```
>>> p_BC_B = frame_B.Pvector(np.r_[3000, 2000, 100].reshape((-1, 1)))
```

**Step 3: Decompose delta BC in E**

```
>>> p_BC_E = p_BC_B.to_ecef_vector()
```

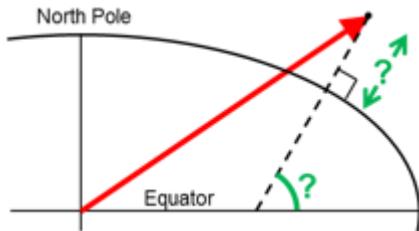
**Step 4: Find point C by adding delta BC to EB**

```
>>> p_EB_E = n_EB_E.to_ecef_vector()
>>> p_EC_E = p_EB_E + p_BC_E
>>> pointC = p_EC_E.to_geo_point()
```

```
>>> lat, lon, z = pointC.latlon_deg
>>> msg = 'Ex2: PosC: lat, lon = {:4.2f}, {:4.2f} deg, height = {:4.2f} m'
>>> msg.format(lat[0], lon[0], -z[0])
'Ex2: PosC: lat, lon = 53.33, 63.47 deg, height = 406.01 m'
```

See also [Example 2 at www.navlab.net](http://www.navlab.net)

### 1.7.3 Example 3: “ECEF-vector to geodetic latitude”



Position B is given as an “ECEF-vector”  $p_{EB\_E}$  (i.e. a vector from E, the center of the Earth, to B, decomposed in E). Find the geodetic latitude, longitude and height ( $\text{lat}_{EB}$ ,  $\text{lon}_{EB}$  and  $h_{EB}$ ), assuming WGS-84 ellipsoid.

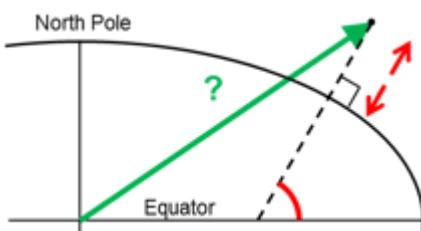
**Solution:**

```
>>> import numpy as np
>>> import nvector as nv
>>> wgs84 = nv.FrameE(name='WGS84')
>>> position_B = 6371e3 * np.vstack((0.9, -1, 1.1)) # m
>>> p_EB_E = wgs84.ECEFvector(position_B)
>>> pointB = p_EB_E.to_geo_point()
```

```
>>> lat, lon, z = pointB.latlon_deg
>>> msg = 'Ex3: Pos B: lat, lon = {:4.2f}, {:4.2f} deg, height = {:9.2f} m'
>>> msg.format(lat[0], lon[0], -z[0])
'Ex3: Pos B: lat, lon = 39.38, -48.01 deg, height = 4702059.83 m'
```

See also [Example 3 at www.navlab.net](http://www.navlab.net)

### 1.7.4 Example 4: “Geodetic latitude to ECEF-vector”



Geodetic latitude, longitude and height are given for position B as  $\text{lat}_{EB}$ ,  $\text{lon}_{EB}$  and  $h_{EB}$ , find the ECEF-vector for this position,  $p_{EB\_E}$ .

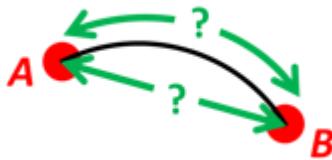
**Solution:**

```
>>> import nvector as nv
>>> wgs84 = nv.FrameE(name='WGS84')
>>> pointB = wgs84.GeoPoint(latitude=1, longitude=2, z=-3, degrees=True)
>>> p_EB_E = pointB.to_ecef_vector()
```

```
>>> 'Ex4: p_EB_E = {} m'.format(p_EB_E.pvector.ravel().tolist())
'Ex4: p_EB_E = [6373290.277218279, 222560.20067473652, 110568.82718178593] m'
```

See also Example 4 at [www.navlab.net](http://www.navlab.net)

### 1.7.5 Example 5: “Surface distance”



Find the surface distance  $s_{AB}$  (i.e. great circle distance) between two positions A and B. The heights of A and B are ignored, i.e. if they don't have zero height, we seek the distance between the points that are at the surface of the Earth, directly above/below A and B. The Euclidean distance (chord length)  $d_{AB}$  should also be found. Use Earth radius 6371e3 m. Compare the results with exact calculations for the WGS-84 ellipsoid.

**Solution for a sphere:**

```
>>> import numpy as np
>>> import nvector as nv
>>> frame_E = nv.FrameE(a=6371e3, f=0)
>>> positionA = frame_E.GeoPoint(latitude=88, longitude=0, degrees=True)
>>> positionB = frame_E.GeoPoint(latitude=89, longitude=-170, degrees=True)
```

```
>>> s_AB, _azia, _azib = positionA.distance_and_azimuth(positionB)
>>> p_AB_E = positionB.to_ecef_vector() - positionA.to_ecef_vector()
>>> d_AB = p_AB_E.length[0]
```

```
>>> msg = 'Ex5: Great circle and Euclidean distance = {}'
>>> msg = msg.format('{:5.2f} km, {:5.2f} km')
>>> msg.format(s_AB / 1000, d_AB / 1000)
'Ex5: Great circle and Euclidean distance = 332.46 km, 332.42 km'
```

**Alternative sphere solution:**

```
>>> path = nv.GeoPath(positionA, positionB)
>>> s_AB2 = path.track_distance(method='greatcircle').ravel()
>>> d_AB2 = path.track_distance(method='euclidean').ravel()
>>> msg.format(s_AB2[0] / 1000, d_AB2[0] / 1000)
'Ex5: Great circle and Euclidean distance = 332.46 km, 332.42 km'
```

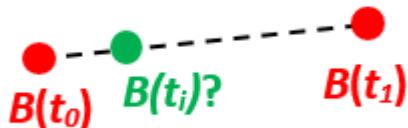
**Exact solution for the WGS84 ellipsoid:**

```
>>> wgs84 = nv.FrameE(name='WGS84')
>>> point1 = wgs84.GeoPoint(latitude=88, longitude=0, degrees=True)
>>> point2 = wgs84.GeoPoint(latitude=89, longitude=-170, degrees=True)
>>> s_12, _azil, _azi2 = point1.distance_and_azimuth(point2)
```

```
>>> p_12_E = point2.to_ecef_vector() - point1.to_ecef_vector()
>>> d_12 = p_12_E.length[0]
>>> msg = 'Ellipsoidal and Euclidean distance = {:.5.2f} km, {:.5.2f} km'
>>> msg.format(s_12 / 1000, d_12 / 1000)
'Ellipsoidal and Euclidean distance = 333.95 km, 333.91 km'
```

See also Example 5 at [www.navlab.net](http://www.navlab.net)

### 1.7.6 Example 6 “Interpolated position”



Given the position of B at time  $t_0$  and  $t_1$ ,  $n_{EB\_E}(t_0)$  and  $n_{EB\_E}(t_1)$ .

Find an interpolated position at time  $t_i$ ,  $n_{EB\_E}(t_i)$ . All positions are given as n-vectors.

**Solution:**

```
>>> import nvector as nv
>>> wgs84 = nv.FrameE(name='WGS84')
>>> n_EB_E_t0 = wgs84.GeoPoint(89, 0, degrees=True).to_nvector()
>>> n_EB_E_t1 = wgs84.GeoPoint(89, 180, degrees=True).to_nvector()
>>> path = nv.GeoPath(n_EB_E_t0, n_EB_E_t1)

>>> t0 = 10.
>>> t1 = 20.
>>> ti = 16. # time of interpolation
>>> ti_n = (ti - t0) / (t1 - t0) # normalized time of interpolation

>>> g_EB_E_ti = path.interpolate(ti_n).to_geo_point()

>>> lat_ti, lon_ti, z_ti = g_EB_E_ti.latlon_deg
>>> msg = 'Ex6, Interpolated position: lat, lon = {:.2.1f} deg, {:.2.1f} deg'
>>> msg.format(lat_ti[0], lon_ti[0])
'Ex6, Interpolated position: lat, lon = 89.8 deg, 180.0 deg'
```

See also Example 6 at [www.navlab.net](http://www.navlab.net)

### 1.7.7 Example 7: “Mean position”

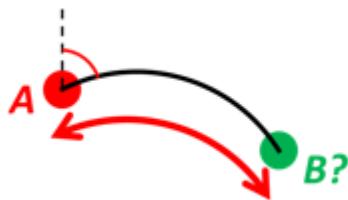


Three positions A, B, and C are given as n-vectors  $n_{EA\_E}$ ,  $n_{EB\_E}$ , and  $n_{EC\_E}$ . Find the mean position, M, given as  $n_{EM\_E}$ . Note that the calculation is independent of the depths of the positions.

**Solution:**

```
>>> import nvector as nv
>>> points = nv.GeoPoint(latitude=[90, 60, 50],
...                         longitude=[0, 10, -20], degrees=True)
>>> nvectors = points.to_nvector()
>>> n_EM_E = nvectors.mean()
>>> g_EM_E = n_EM_E.to_geo_point()
>>> lat, lon = g_EM_E.latitude_deg, g_EM_E.longitude_deg
>>> msg = 'Ex7: Pos M: lat, lon = {:.4.2f}, {:.4.2f} deg'
>>> msg.format(lat[0], lon[0])
'Ex7: Pos M: lat, lon = 67.24, -6.92 deg'
```

See also Example 7 at [www.navlab.net](http://www.navlab.net)

**1.7.8 Example 8: “A and azimuth/distance to B”**

We have an initial position A, direction of travel given as an azimuth (bearing) relative to north (clockwise), and finally the distance to travel along a great circle given as sAB. Use Earth radius 6371e3 m to find the destination point B.

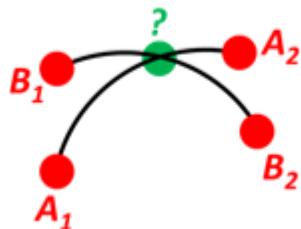
In geodesy this is known as “The first geodetic problem” or “The direct geodetic problem” for a sphere, and we see that this is similar to Example 2, but now the delta is given as an azimuth and a great circle distance. (“The second/inverse geodetic problem” for a sphere is already solved in Examples 1 and 5.)

**Solution:**

```
>>> import nvector as nv
>>> frame = nv.FrameE(a=6371e3, f=0)
>>> pointA = frame.GeoPoint(latitude=80, longitude=-90, degrees=True)
>>> pointB, _azimuthb = pointA.displace(distance=1000, azimuth=200,
...                                         degrees=True)
...                                         degrees=True)
>>> lat, lon = pointB.latitude_deg, pointB.longitude_deg

>>> msg = 'Ex8, Destination: lat, lon = {:.4.2f} deg, {:.4.2f} deg'
>>> msg.format(lat, lon)
'Ex8, Destination: lat, lon = 79.99 deg, -90.02 deg'
```

See also Example 8 at [www.navlab.net](http://www.navlab.net)

**1.7.9 Example 9: “Intersection of two paths”**

Define a path from two given positions (at the surface of a spherical Earth), as the great circle that goes through the two points.

Path A is given by A1 and A2, while path B is given by B1 and B2.

Find the position C where the two great circles intersect.

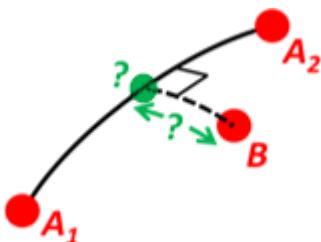
**Solution:**

```
>>> import nvector as nv
>>> pointA1 = nv.GeoPoint(10, 20, degrees=True)
>>> pointA2 = nv.GeoPoint(30, 40, degrees=True)
>>> pointB1 = nv.GeoPoint(50, 60, degrees=True)
>>> pointB2 = nv.GeoPoint(70, 80, degrees=True)
>>> pathA = nv.GeoPath(pointA1, pointA2)
>>> pathB = nv.GeoPath(pointB1, pointB2)
```

```
>>> pointC = pathA.intersect(pathB)
>>> np.allclose(pathA.on_path(pointC), pathB.on_path(pointC))
True
>>> np.allclose(pathA.on_great_circle(pointC),
...                 pathB.on_great_circle(pointC))
True
>>> pointC = pointC.to_geo_point()
>>> lat, lon = pointC.latitude_deg, pointC.longitude_deg
>>> msg = 'Ex9, Intersection: lat, lon = {:.4f}, {:.4f} deg'
>>> msg.format(lat[0], lon[0])
'Ex9, Intersection: lat, lon = 40.32, 55.90 deg'
```

See also Example 9 at [www.navlab.net](http://www.navlab.net)

### 1.7.10 Example 10: “Cross track distance”



Path A is given by the two positions A1 and A2 (similar to the previous example).

Find the cross track distance  $s_{xt}$  between the path A (i.e. the great circle through A1 and A2) and the position B (i.e. the shortest distance at the surface, between the great circle and B).

Also find the Euclidean distance  $d_{xt}$  between B and the plane defined by the great circle. Use Earth radius  $6371\text{e}3$ .

Finally, find the intersection point on the great circle and determine if it is between position A1 and A2.

**Solution:**

```
>>> import nvector as nv
>>> frame = nv.FrameE(a=6371e3, f=0)
>>> pointA1 = frame.GeoPoint(0, 0, degrees=True)
>>> pointA2 = frame.GeoPoint(10, 0, degrees=True)
>>> pointB = frame.GeoPoint(1, 0.1, degrees=True)
>>> pathA = nv.GeoPath(pointA1, pointA2)
```

```
>>> s_xt = pathA.cross_track_distance(pointB, method='greatcircle').ravel()
>>> d_xt = pathA.cross_track_distance(pointB, method='euclidean').ravel()
```

```
>>> val_txt = '{:4.2f} km, {:4.2f} km'.format(s_xt[0]/1000, d_xt[0]/1000)
>>> 'Ex10: Cross track distance: s_xt, d_xt = {}'.format(val_txt)
'Ex10: Cross track distance: s_xt, d_xt = 11.12 km, 11.12 km'
```

```
>>> pointC = pathA.closest_point_on_great_circle(pointB)
>>> np.allclose(pathA.on_path(pointC), True)
True
```

See also Example 10 at [www.navlab.net](http://www.navlab.net)

## 1.8 See also

[geographiclib](#)

## 1.9 Functional examples

Below the functional solution to some common geodesic problems are given. In the first example the object-oriented solution is also given. The object-oriented solutions to the remaining problems can be found [here](#).

### 1.9.1 Example 1: “A and B to delta”



Given two positions, A and B as latitudes, longitudes and depths relative to Earth, E.

Find the exact vector between the two positions, given in meters north, east, and down, and find the direction (azimuth) to B, relative to north. Assume WGS-84 ellipsoid. The given depths are from the ellipsoid surface. Use position A to define north, east, and down directions. (Due to the curvature of Earth and different directions to the North Pole, the north, east, and down directions will change (relative to Earth) for different places. A must be outside the poles for the north and east directions to be defined.)

**Solution:**

```
>>> import numpy as np
>>> import nvector as nv
>>> from nvector import rad, deg
```

```
>>> lat_EA, lon_EA, z_EA = rad(1), rad(2), 3
>>> lat_EB, lon_EB, z_EB = rad(4), rad(5), 6
```

**Step1: Convert to n-vectors:**

```
>>> n_EA_E = nv.lat_lon2n_E(lat_EA, lon_EA)
>>> n_EB_E = nv.lat_lon2n_E(lat_EB, lon_EB)
```

**Step2: Find p\_AB\_E (delta decomposed in E).WGS-84 ellipsoid is default:**

```
>>> p_AB_E = nv.n_EA_E_and_n_EB_E2p_AB_E(n_EA_E, n_EB_E, z_EA, z_EB)
```

**Step3: Find R\_EN for position A:**

```
>>> R_EN = nv.n_E2R_EN(n_EA_E)
```

**Step4: Find p\_AB\_N (delta decomposed in N).**

```
>>> p_AB_N = np.dot(R_EN.T, p_AB_E).ravel()
>>> valtxt = '{0:8.2f}, {1:8.2f}, {2:8.2f}'.format(*p_AB_N)
>>> 'Ex1: delta north, east, down = {}'.format(valtxt)
'Ex1: delta north, east, down = 331730.23, 332997.87, 17404.27'
```

**Step5: Also find the direction (azimuth) to B, relative to north:**

```
>>> azimuth = np.arctan2(p_AB_N[1], p_AB_N[0])
>>> 'azimuth = {0:4.2f} deg'.format(deg(azimuth))
'azimuth = 45.11 deg'
```

**OO-Solution:**

```
>>> import numpy as np
>>> import nvector as nv
>>> wgs84 = nv.FrameE(name='WGS84')
>>> pointA = wgs84.GeoPoint(latitude=1, longitude=2, z=3, degrees=True)
>>> pointB = wgs84.GeoPoint(latitude=4, longitude=5, z=6, degrees=True)
```

**Step1: Find p\_AB\_N (delta decomposed in N).**

```
>>> p_AB_N = pointA.delta_to(pointB)
>>> x, y, z = p_AB_N.pvector.ravel()
>>> valtxt = '{0:8.2f}, {1:8.2f}, {2:8.2f}'.format(x, y, z)
>>> 'Ex1: delta north, east, down = {}'.format(valtxt)
'Ex1: delta north, east, down = 331730.23, 332997.87, 17404.27'
```

**Step2: Also find the direction (azimuth) to B, relative to north:**

```
>>> azimuth = p_AB_N.azimuth_deg[0]
>>> 'azimuth = {0:4.2f} deg'.format(azimuth)
'azimuth = 45.11 deg'
```

See also Example 1 at [www.navlab.net](http://www.navlab.net)

**1.9.2 Example 2: “B and delta to C”**

A radar or sonar attached to a vehicle B (Body coordinate frame) measures the distance and direction to an object C. We assume that the distance and two angles (typically bearing and elevation relative to B) are already combined to the vector p\_BC\_B (i.e. the vector from B to C, decomposed in B). The position of B is given as n\_EB\_E and z\_EB, and the orientation (attitude) of B is given as R\_NB (this rotation matrix can be found from roll/pitch/yaw by using zyxR).

Find the exact position of object C as n-vector and depth ( n\_EC\_E and z\_EC ), assuming Earth ellipsoid with semi-major axis a and flattening f. For WGS-72, use a = 6 378 135 m and f = 1/298.26.

**Solution:**

```
>>> import numpy as np
>>> import nvector as nv
>>> from nvector import rad, deg
```

A custom reference ellipsoid is given (replacing WGS-84):

```
>>> wgs72 = dict(a=6378135, f=1.0/298.26)
```

**Step 1 Position and orientation of B is 400m above E:**

```
>>> n_EB_E = nv.unit([[1], [2], [3]]) # unit to get unit length of vector
>>> z_EB = -400
>>> yaw, pitch, roll = rad(10), rad(20), rad(30)
>>> R_NB = nv.zyx2R(yaw, pitch, roll)
```

**Step 2: Delta BC decomposed in B**

```
>>> p_BC_B = np.r_[3000, 2000, 100].reshape((-1, 1))
```

**Step 3: Find R\_EN:**

```
>>> R_EN = nv.n_E2R_EN(n_EB_E)
```

**Step 4: Find R\_EB, from R\_EN and R\_NB:**

```
>>> R_EB = np.dot(R_EN, R_NB) # Note: closest frames cancel
```

**Step 5: Decompose the delta BC vector in E:**

```
>>> p_BC_E = np.dot(R_EB, p_BC_B)
```

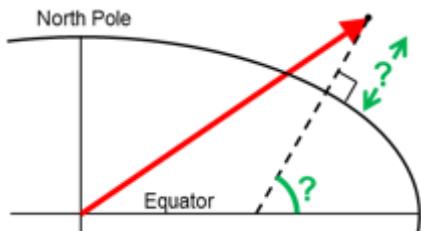
**Step 6: Find the position of C, using the functions that goes from one**

```
>>> n_EC_E, z_EC = nv.n_EA_E_and_p_AB_E2n_EB_E(n_EB_E, p_BC_E, z_EB, **wgs72)
```

```
>>> lat_EC, lon_EC = nv.n_E2lat_lon(n_EC_E)
>>> lat, lon, z = deg(lat_EC), deg(lon_EC), z_EC
>>> msg = 'Ex2: PosC: lat, lon = {:.4.2f}, {:.4.2f} deg, height = {:.4.2f} m'
>>> msg.format(lat[0], lon[0], -z[0])
'Ex2: PosC: lat, lon = 53.33, 63.47 deg, height = 406.01 m'
```

See also Example 2 at [www.navlab.net](http://www.navlab.net)

### 1.9.3 Example 3: “ECEF-vector to geodetic latitude”



Position B is given as an “ECEF-vector”  $p_{EB\_E}$  (i.e. a vector from E, the center of the Earth, to B, decomposed in E). Find the geodetic latitude, longitude and height ( $\text{lat}_{EB}$ ,  $\text{lon}_{EB}$  and  $h_{EB}$ ), assuming WGS-84 ellipsoid.

**Solution:**

```

>>> import numpy as np
>>> import nvector as nv
>>> from nvector import deg
>>> wgs84 = dict(a=6378137.0, f=1.0/298.257223563)
>>> p_EB_E = 6371e3 * np.vstack((0.9, -1, 1.1)) # m

>>> n_EB_E, z_EB = nv.p_EB_E2n_EB_E(p_EB_E, **wgs84)

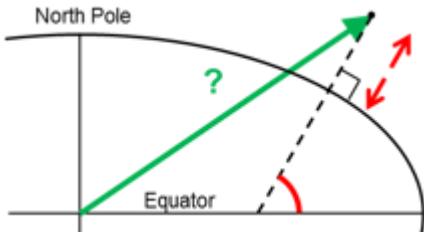
>>> lat_EB, lon_EB = nv.n_E2lat_lon(n_EB_E)
>>> h = -z_EB
>>> lat, lon = deg(lat_EB), deg(lon_EB)

>>> msg = 'Ex3: Pos B: lat, lon = {:.4f}, {:.4f} deg, height = {:.9f} m'
>>> msg.format(lat[0], lon[0], h[0])
'Ex3: Pos B: lat, lon = 39.38, -48.01 deg, height = 4702059.83 m'

```

See also Example 3 at [www.navlab.net](http://www.navlab.net)

#### 1.9.4 Example 4: “Geodetic latitude to ECEF-vector”



Geodetic latitude, longitude and height are given for position B as latEB, lonEB and hEB, find the ECEF-vector for this position, p\_EB\_E.

**Solution:**

```

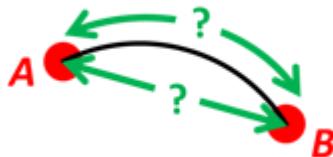
>>> import nvector as nv
>>> from nvector import rad
>>> wgs84 = dict(a=6378137.0, f=1.0/298.257223563)
>>> lat_EB, lon_EB = rad(1), rad(2)
>>> h_EB = 3
>>> n_EB_E = nv.lat_lon2n_E(lat_EB, lon_EB)
>>> p_EB_E = nv.n_EB_E2p_EB_E(n_EB_E, -h_EB, **wgs84)

>>> 'Ex4: p_EB_E = {} m'.format(p_EB_E.ravel().tolist())
'Ex4: p_EB_E = [6373290.277218279, 222560.20067473652, 110568.82718178593] m'

```

See also Example 4 at [www.navlab.net](http://www.navlab.net)

#### 1.9.5 Example 5: “Surface distance”



Find the surface distance  $s_{AB}$  (i.e. great circle distance) between two positions A and B. The heights of A and B are ignored, i.e. if they don't have zero height, we seek the distance between the points that are at the surface of the Earth, directly above/below A and B. The Euclidean distance (chord length)  $d_{AB}$  should also be found. Use Earth radius 6371e3 m. Compare the results with exact calculations for the WGS-84 ellipsoid.

**Solution for a sphere:**

```
>>> import numpy as np
>>> import nvector as nv
>>> from nvector import rad

>>> n_EA_E = nv.lat_lon2n_E(rad(88), rad(0))
>>> n_EB_E = nv.lat_lon2n_E(rad(89), rad(-170))

>>> r_Earth = 6371e3 # m, mean Earth radius
>>> s_AB = nv.great_circle_distance(n_EA_E, n_EB_E, radius=r_Earth)[0]
>>> d_AB = nv.euclidean_distance(n_EA_E, n_EB_E, radius=r_Earth)[0]

>>> msg = 'Ex5: Great circle and Euclidean distance = {}'
>>> msg = msg.format('{:5.2f} km, {:5.2f} km')
>>> msg.format(s_AB / 1000, d_AB / 1000)
'Ex5: Great circle and Euclidean distance = 332.46 km, 332.42 km'
```

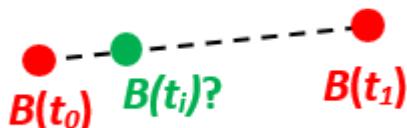
**Exact solution for the WGS84 ellipsoid:**

```
>>> wgs84 = nv.FrameE(name='WGS84')
>>> point1 = wgs84.GeoPoint(latitude=88, longitude=0, degrees=True)
>>> point2 = wgs84.GeoPoint(latitude=89, longitude=-170, degrees=True)
>>> s_12, _azil, _azi2 = point1.distance_and_azimuth(point2)

>>> p_12_E = point2.to_ecef_vector() - point1.to_ecef_vector()
>>> d_12 = p_12_E.length[0]
>>> msg = 'Ellipsoidal and Euclidean distance = {:5.2f} km, {:5.2f} km'
>>> msg.format(s_12 / 1000, d_12 / 1000)
'Ellipsoidal and Euclidean distance = 333.95 km, 333.91 km'
```

See also Example 5 at [www.navlab.net](http://www.navlab.net)

### 1.9.6 Example 6 “Interpolated position”



Given the position of B at time  $t_0$  and  $t_1$ ,  $n_{EB\_E}(t_0)$  and  $n_{EB\_E}(t_1)$ .

Find an interpolated position at time  $t_i$ ,  $n_{EB\_E}(t_i)$ . All positions are given as n-vectors.

**Solution:**

```
>>> import nvector as nv
>>> from nvector import rad, deg
>>> n_EB_E_t0 = nv.lat_lon2n_E(rad(89), rad(0))
>>> n_EB_E_t1 = nv.lat_lon2n_E(rad(89), rad(180))
```

```
>>> t0 = 10.
>>> t1 = 20.
>>> ti = 16. # time of interpolation
>>> ti_n = (ti - t0) / (t1 - t0) # normalized time of interpolation
```

```
>>> n_EB_E_ti = nv.unit(n_EB_E_t0 + ti_n * (n_EB_E_t1 - n_EB_E_t0))
>>> lat_EB_ti, lon_EB_ti = nv.n_E2lat_lon(n_EB_E_ti)
```

```
>>> lat_ti, lon_ti = deg(lat_EB_ti), deg(lon_EB_ti)
>>> msg = 'Ex6, Interpolated position: lat, lon = {:2.1f} deg, {:2.1f} deg'
>>> msg.format(lat_ti[0], lon_ti[0])
'Ex6, Interpolated position: lat, lon = 89.8 deg, 180.0 deg'
```

See also Example 6 at [www.navlab.net](http://www.navlab.net)

### 1.9.7 Example 7: “Mean position”



Three positions A, B, and C are given as n-vectors  $n\_EA\_E$ ,  $n\_EB\_E$ , and  $n\_EC\_E$ . Find the mean position, M, given as  $n\_EM\_E$ . Note that the calculation is independent of the depths of the positions.

**Solution:**

```
>>> import numpy as np
>>> import nvector as nv
>>> from nvector import rad, deg
```

```
>>> n_EA_E = nv.lat_lon2n_E(rad(90), rad(0))
>>> n_EB_E = nv.lat_lon2n_E(rad(60), rad(10))
>>> n_EC_E = nv.lat_lon2n_E(rad(50), rad(-20))
```

```
>>> n_EM_E = nv.unit(n_EA_E + n_EB_E + n_EC_E)
```

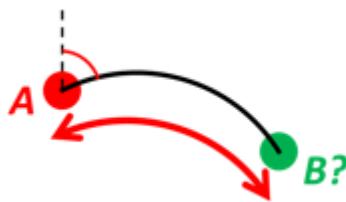
or

```
>>> n_EM_E = nv.mean_horizontal_position(np.hstack((n_EA_E, n_EB_E, n_EC_E)))
```

```
>>> lat, lon = nv.n_E2lat_lon(n_EM_E)
>>> lat, lon = deg(lat), deg(lon)
>>> msg = 'Ex7: Pos M: lat, lon = {:4.2f}, {:4.2f} deg'
>>> msg.format(lat[0], lon[0])
'Ex7: Pos M: lat, lon = 67.24, -6.92 deg'
```

See also Example 7 at [www.navlab.net](http://www.navlab.net)

### 1.9.8 Example 8: “A and azimuth/distance to B”



We have an initial position A, direction of travel given as an azimuth (bearing) relative to north (clockwise), and finally the distance to travel along a great circle given as  $s_{AB}$ . Use Earth radius  $6371\text{e}3$  m to find the destination point B.

In geodesy this is known as “The first geodetic problem” or “The direct geodetic problem” for a sphere, and we see that this is similar to [Example 2](#), but now the delta is given as an azimuth and a great circle distance. (“The second/inverse geodetic problem” for a sphere is already solved in Examples [1](#) and [5](#).)

**Solution:**

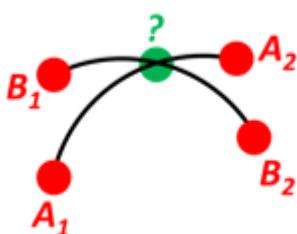
```
>>> import nvector as nv
>>> from nvector import rad, deg
>>> lat, lon = rad(80), rad(-90)

>>> n_EA_E = nv.lat_lon2n_E(lat, lon)
>>> azimuth = rad(200)
>>> s_AB = 1000.0 # [m]
>>> r_earth = 6371e3 # [m], mean earth radius

>>> distance_rad = s_AB / r_earth
>>> n_EB_E = nv.n_EA_E_distance_and_azimuth2n_EB_E(n_EA_E, distance_rad,
...                                         azimuth)
>>> lat_EB, lon_EB = nv.n_E2lat_lon(n_EB_E)
>>> lat, lon = deg(lat_EB), deg(lon_EB)
>>> msg = 'Ex8, Destination: lat, lon = {:.4f} deg, {:.4f} deg'
>>> msg.format(lat[0], lon[0])
'Ex8, Destination: lat, lon = 79.99 deg, -90.02 deg'
```

See also [Example 8](#) at [www.navlab.net](http://www.navlab.net)

### 1.9.9 Example 9: “Intersection of two paths”



Define a path from two given positions (at the surface of a spherical Earth), as the great circle that goes through the two points.

Path A is given by A<sub>1</sub> and A<sub>2</sub>, while path B is given by B<sub>1</sub> and B<sub>2</sub>.

Find the position C where the two great circles intersect.

**Solution:**

```
>>> import numpy as np
>>> import nvector as nv
>>> from nvector import rad, deg
```

```
>>> n_EA1_E = nv.lat_lon2n_E(rad(10), rad(20))
>>> n_EA2_E = nv.lat_lon2n_E(rad(30), rad(40))
>>> n_EB1_E = nv.lat_lon2n_E(rad(50), rad(60))
>>> n_EB2_E = nv.lat_lon2n_E(rad(70), rad(80))
```

```
>>> n_EC_E = nv.unit(np.cross(np.cross(n_EA1_E, n_EA2_E, axis=0),
...                             np.cross(n_EB1_E, n_EB2_E, axis=0),
...                             axis=0))
>>> n_EC_E *= np.sign(np.dot(n_EC_E.T, n_EA1_E))
```

or alternatively

```
>>> path_a, path_b = (n_EA1_E, n_EA2_E), (n_EB1_E, n_EB2_E)
>>> n_EC_E = nv.intersect(path_a, path_b)
```

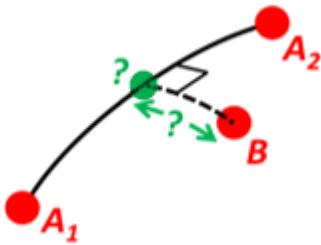
```
>>> lat_EC, lon_EC = nv.n_E2lat_lon(n_EC_E)
```

```
>>> lat, lon = deg(lat_EC), deg(lon_EC)
>>> msg = 'Ex9, Intersection: lat, lon = {:.2f}, {:.2f} deg'
>>> msg.format(lat[0], lon[0])
'Ex9, Intersection: lat, lon = 40.32, 55.90 deg'
```

```
>>> np.allclose(nv.on_great_circle_path(path_a, n_EC_E),
...               nv.on_great_circle_path(path_b, n_EC_E))
True
>>> np.allclose(nv.on_great_circle(path_a, n_EC_E), nv.on_great_circle(path_b, n_EC_E))
True
```

See also Example 9 at [www.navlab.net](http://www.navlab.net)

### 1.9.10 Example 10: “Cross track distance”



Path A is given by the two positions A1 and A2 (similar to the previous example).

Find the cross track distance  $sxt$  between the path A (i.e. the great circle through A1 and A2) and the position B (i.e. the shortest distance at the surface, between the great circle and B).

Also find the Euclidean distance  $dxt$  between B and the plane defined by the great circle. Use Earth radius  $6371e3$ .

Finally, find the intersection point on the great circle and determine if it is between position A1 and A2.

**Solution:**

```
>>> import numpy as np
>>> import nvector as nv
>>> n_EA1_E = nv.lat_lon2n_E(rad(0), rad(0))
```

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```
>>> n_EA2_E = nv.lat_lon2n_E(rad(10), rad(0))
>>> n_EB_E = nv.lat_lon2n_E(rad(1), rad(0.1))
>>> path = (n_EA1_E, n_EA2_E)
>>> radius = 6371e3 # mean earth radius [m]
>>> s_xt = nv.cross_track_distance(path, n_EB_E, radius=radius)
>>> d_xt = nv.cross_track_distance(path, n_EB_E, method='euclidean',
...                                     radius=radius)
```

```
>>> val_txt = '{:4.2f} km, {:4.2f} km'.format(s_xt[0]/1000, d_xt[0]/1000)
>>> 'Ex10: Cross track distance: s_xt, d_xt = {}'.format(val_txt)
'Ex10: Cross track distance: s_xt, d_xt = 11.12 km, 11.12 km'
```

```
>>> n_EC_E = nv.closest_point_on_great_circle(path, n_EB_E)
>>> np.allclose(nv.on_great_circle_path(path, n_EC_E, radius), True)
True
```

**Alternative solution 2:**

```
>>> s_xt2 = nv.great_circle_distance(n_EB_E, n_EC_E, radius)
>>> d_xt2 = nv.euclidean_distance(n_EB_E, n_EC_E, radius)
>>> np.allclose(s_xt, s_xt2), np.allclose(d_xt, d_xt2)
(True, True)
```

**Alternative solution 3:**

```
>>> c_E = nv.great_circle_normal(n_EA1_E, n_EA2_E)
>>> sin_theta = -np.dot(c_E.T, n_EB_E).ravel()
>>> s_xt3 = np.arcsin(sin_theta) * radius
>>> d_xt3 = sin_theta * radius
>>> np.allclose(s_xt, s_xt3), np.allclose(d_xt, d_xt3)
(True, True)
```

See also Example 10 at [www.navlab.net](http://www.navlab.net)

## 1.10 License

The content of this library **is** based on the following publication:

Gade, K. (2010). A Nonsingular Horizontal Position Representation, The Journal of Navigation, Volume 63, Issue 03, pp 395-417, July 2010.  
[\(www.navlab.net/Publications/A\\_Nonsingular\\_Horizontal\\_Position\\_Representation.pdf\)](http://www.navlab.net/Publications/A_Nonsingular_Horizontal_Position_Representation.pdf)

This paper should be cited **in** publications using this library.

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## 1.11 Developers

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- Per A. Brodtkorb FFI

## 1.12 Changelog

### 1.12.1 Version 0.6.0, December 09, 2018

#### Per A Brodtkorb (79):

- Updated requirements in setup.py
- Removed tox.ini
- Updated documentation on how to set package version
- Made a separate script to set package version in nvector/\_\_init\_\_.py
- Updated docstring for select\_ellipsoid
- Replace GeoPoint.geo\_point with GeoPoint.displace and removed deprecated GeoPoint.geo\_point
- Update .travis.yml
- Fix so that codeclimate is able to parse .travis.yml
- Only run sonar and codeclimate reporter for python v3.6
- Added sonar-project.properties
- **Pinned coverage to v4.3.4 due to fact that codeclimate reporter is only compatible with Coverage.py versions >=4.0,<4.4.**
- Updated with sonar scanner.
- Added .pylintrc
- Set up codeclimate reporter
- Updated docstring for unit function.
- Avoid division by zero in unit function.
- Reenabled the doctest of plot\_mean\_position

- Reset “pyscaffold==2.5.11”
- Replaced deprecated basemap with cartopy.
- **Replaced doctest of plot\_mean\_position with test\_plot\_mean\_position in test\_plot.py**
- **Fixed failing doctests for python v3.4 and v3.5 and made them more robust.**
- Fixed failing doctests and made them more robust.
- Increased pycoverage version to use.
- moved nvector to src/nvector/
- **Reset the setup.py to require ‘pyscaffold==2.5.11’ which works on python version 3.4, 3.5 and 3.6. as well as 2.7**
- Updated unittests.
- Updated tests.
- Removed obsolete code
- Added test for delta\_L
- **Added corner testcase for pointA.displace(distance=1000,azimuth=np.deg2rad(200))**
- Added test for path.track\_distance(method='exact')
- **Added delta\_L a function that return cartesian delta vector from positions A to B decomposed in L.**
- Simplified OO-solution in example 1 by using delta\_N function
- Refactored duplicated code
- **Vectorized code so that the frames can take more than one position at the time.**
- Keeping only the html docs in the distribution.
- **replaced link from latest to stable docs on readthedocs and updated crosstrack distance test.**
- updated documentation in setup.py

### 1.12.2 Version 0.5.2, Mars 7, 2017

#### Per A Brodkorb (10):

- Fixed tests in tests/test\_frames.py
- Updated to setup.cfg and tox.ini + pep8
- updated .travis.yml
- Updated Readme.rst with new example 10 picture and link to nvector docs at readthedocs.
- updated official documentation links
- Updated crosstrack distance tests.

### 1.12.3 Version 0.5.1, Mars 5, 2017

#### Cody (4):

- Explicitly numbered replacement fields
- Migrated % string formating

#### Per A Brodkorb (29):

- pep8

- Updated failing examples
- Updated README.rst
- Removed obsolete pass statement
- Documented functions
- added .checkignore for quantifycode
- moved test\_docstrings and use\_docstring\_from into \_common.py
- Added .codeclimate.yml
- Updated installation information in \_info.py
- Added GeoPath.on\_path method. Clarified intersection example
- Added great\_circle\_normal, cross\_track\_distance Renamed intersection to intersect (Intersection is deprecated.)
- Simplified R2zyx with a call to R2xyz Improved accuracy for great circle cross track distance for small distances.
- Added on\_great\_circle, \_on\_great\_circle\_path, \_on\_ellipsoid\_path, closest\_point\_on\_great\_circle and closest\_point\_on\_path to GeoPath
- made \_\_eq\_\_ more robust for frames
- Removed duplicated code
- Updated tests
- Removed fishy test
- replaced zero n-vector with nan
- Commented out failing test.
- Added example 10 image
- Added ‘closest\_point\_on\_great\_circle’, ‘on\_great\_circle’, ‘on\_great\_circle\_path’.
- Updated examples + documentation
- Updated index depth
- Updated README.rst and classifier in setup.cfg

#### 1.12.4 Version 0.4.1, Januar 19, 2016

pbrod (46):

- Cosmetic updates
- Updated README.rst
- updated docs and removed unused code
- updated README.rst and .coveragerc
- Refactored out \_check\_frames
- Refactored out \_default\_frame
- Updated .coveragerc
- Added link to geographiclib
- Updated external link
- Updated documentation
- Added figures to examples

- Added GeoPath.interpolate + interpolation example 6
- Added links to FFI homepage.
- **Updated documentation:**
  - Added link to nvector toolbox for matlab
  - For each example added links to the more detailed explanation on the homepage
- Updated link to nvector toolbox for matlab
- Added link to nvector on pypi
- Updated documentation fro FrameB, FrameE, FrameL and FrameN.
- updated \_\_all\_\_ variable
- Added missing R\_Ee to function n\_EA\_E\_and\_n\_EB\_E2azimuth + updated documentation
- Updated CHANGES.rst
- Updated conf.py
- Renamed info.py to \_info.py
- All examples are now generated from \_examples.py.

## 1.12.5 Version 0.1.3, Januar 1, 2016

pbrod (31):

- Refactored
- Updated tests
- Updated docs
- Moved tests to nvector/tests
- Updated .coverage Added travis.yml, .landscape.yml
- Deleted obsolete LICENSE
- Updated README.rst
- Removed ngs version
- Fixed bug in .travis.yml
- Updated .travis.yml
- Removed dependence on navigator.py
- Updated README.rst
- Updated examples
- Deleted skeleton.py and added tox.ini
- Small refactoring Renamed distance\_rad\_bearing\_rad2point to n\_EA\_E\_distance\_and\_azimuth2n\_EB\_E updated tests
- Renamed azimuth to n\_EA\_E\_and\_n\_EB\_E2azimuth Added tests for R2xyz as well as R2zyx
- Removed backward compatibility Added test\_n\_E\_and\_wa2R\_EL
- Refactored tests
- Commented out failing tests on python 3+
- updated CHANGES.rst
- Removed bug in setup.py

## 1.12.6 Version 0.1.1, Januar 1, 2016

**pbrod (31):**

- Initial commit: Translated code from Matlab to Python.
- Added object oriented interface to nvector library
- Added tests for object oriented interface
- Added geodesic tests.

## 1.13 Modules

**Release 0.6**

**Date** Dec 09, 2018

This reference manual details functions, modules, and objects included in nvector, describing what they are and what they do.

### 1.13.1 nvector package

#### 1.13.1.1 Geodesic functions

<code>lat_lon2n_E(latitude, longitude[, R_Ee])</code>	Converts latitude and longitude to n-vector.
<code>n_E2lat_lon(n_E[, R_Ee])</code>	Converts n-vector to latitude and longitude.
<code>n_EB_E2p_EB_E(n_EB_E[, depth, a, f, R_Ee])</code>	Converts n-vector to Cartesian position vector in meters.
<code>p_EB_E2n_EB_E(p_EB_E[, a, f, R_Ee])</code>	Converts Cartesian position vector in meters to n-vector.
<code>n_EA_E_and_n_EB_E2p_AB_E(n_EA_E, n_EB_E[, ...])</code>	Return the delta vector from position A to B.
<code>n_EA_E_and_p_AB_E2n_EB_E(n_EA_E, p_AB_E[, ...])</code>	Return position B from position A and delta.
<code>n_EA_E_and_n_EB_E2azimuth(n_EA_E, n_EB_E[, ...])</code>	Return azimuth from A to B, relative to North.
<code>n_EA_E_distance_and_azimuth2n_EB_E(n_EA_E, ...)</code>	Return position B from azimuth and distance from position A
<code>great_circle_distance(n_EA_E, n_EB_E[, radius])</code>	Return great circle distance between positions A and B
<code>euclidean_distance(n_EA_E, n_EB_E[, radius])</code>	Return Euclidean distance between positions A and B
<code>cross_track_distance(path, n_EB_E[, method, ...])</code>	Return cross track distance between path A and position B.
<code>closest_point_on_great_circle(path, n_EB_E)</code>	Return closest point C on great circle path A to position B.
<code>intersect(path_a, path_b)</code>	Return the intersection(s) between the great circles of the two paths
<code>mean_horizontal_position(n_EB_E)</code>	Return the n-vector of the horizontal mean position.
<code>on_great_circle(path, n_EB_E[, radius, ...])</code>	True if position B is on great circle through path A.
<code>on_great_circle_path(path, n_EB_E[, radius, ...])</code>	True if position B is on great circle and between endpoints of path A.

### 1.13.1.1.1 nvector.\_core.lat\_lon2n\_E

**lat\_lon2n\_E** (*latitude, longitude, R\_Ee=None*)

Converts latitude and longitude to n-vector.

#### Parameters

**latitude, longitude: real scalars or vectors of length n.** Geodetic latitude and longitude given in [rad]

**R\_Ee** [3 x 3 array] rotation matrix defining the axes of the coordinate frame E.

#### Returns

**n\_E: 3 x n array** n-vector(s) [no unit] decomposed in E.

See also:

[n\\_E2lat\\_lon](#)

### 1.13.1.1.2 nvector.\_core.n\_E2lat\_lon

**n\_E2lat\_lon** (*n\_E, R\_Ee=None*)

Converts n-vector to latitude and longitude.

#### Parameters

**n\_E: 3 x n array** n-vector [no unit] decomposed in E.

**R\_Ee** [3 x 3 array] rotation matrix defining the axes of the coordinate frame E.

#### Returns

**latitude, longitude: real scalars or vectors of length n.** Geodetic latitude and longitude given in [rad]

See also:

[lat\\_lon2n\\_E](#)

### 1.13.1.1.3 nvector.\_core.n\_EB\_E2p\_EB\_E

**n\_EB\_E2p\_EB\_E** (*n\_EB\_E, depth=0, a=6378137, f=0.0033528106647474805, R\_Ee=None*)

Converts n-vector to Cartesian position vector in meters.

#### Parameters

**n\_EB\_E: 3 x n array**

n-vector(s) [no unit] of position B, decomposed in E.

**depth: 1 x n array** Depth(s) [m] of system B, relative to the ellipsoid (depth = -height)

**a: real scalar, default WGS-84 ellipsoid.** Semi-major axis of the Earth ellipsoid given in [m].

**f: real scalar, default WGS-84 ellipsoid.** Flattening [no unit] of the Earth ellipsoid. If  $f==0$  then spherical Earth with radius a is used instead of WGS-84.

**R\_Ee** [3 x 3 array] rotation matrix defining the axes of the coordinate frame E.

#### Returns

**p\_EB\_E: 3 x n array** Cartesian position vector(s) from E to B, decomposed in E.

#### 1.13.1.1.4 Notes

The position of B (typically body) relative to E (typically Earth) is given into this function as n-vector, n\_EB\_E. The function converts to cartesian position vector (“ECEF-vector”), p\_EB\_E, in meters. The calculation is exact, taking the ellipsity of the Earth into account. It is also non-singular as both n-vector and p-vector are non-singular (except for the center of the Earth). The default ellipsoid model used is WGS-84, but other ellipsoids/spheres might be specified.

#### 1.13.1.1.5 nvector.\_core.p\_EB\_E2n\_EB\_E

**p\_EB\_E2n\_EB\_E** ( $p_{EB\_E}$ ,  $a=6378137$ ,  $f=0.0033528106647474805$ ,  $R_{Ee}=None$ )

Converts Cartesian position vector in meters to n-vector.

##### Parameters

**p\_EB\_E: 3 x n array**

Cartesian position vector(s) from E to B, decomposed in E.

**a: real scalar, default WGS-84 ellipsoid.** Semi-major axis of the Earth ellipsoid given in [m].

**f: real scalar, default WGS-84 ellipsoid.** Flattening [no unit] of the Earth ellipsoid. If  $f==0$  then spherical Earth with radius a is used in stead of WGS-84.

**R\_Ee** [3 x 3 array] rotation matrix defining the axes of the coordinate frame E.

##### Returns

**n\_EB\_E: 3 x n array**

n-vector(s) [no unit] of position B, decomposed in E.

**depth: 1 x n array** Depth(s) [m] of system B, relative to the ellipsoid (depth = -height)

#### 1.13.1.1.6 Notes

The position of B (typically body) relative to E (typically Earth) is given into this function as cartesian position vector p\_EB\_E, in meters. (“ECEF-vector”). The function converts to n-vector, n\_EB\_E and its depth, depth. The calculation is exact, taking the ellipsity of the Earth into account. It is also non-singular as both n-vector and p-vector are non-singular (except for the center of the Earth). The default ellipsoid model used is WGS-84, but other ellipsoids/spheres might be specified.

#### 1.13.1.1.7 nvector.\_core.n\_EA\_E\_and\_n\_EB\_E2p\_AB\_E

**n\_EA\_E\_and\_n\_EB\_E2p\_AB\_E** ( $n_{EA\_E}$ ,  $n_{EB\_E}$ ,  $z_{EA}=0$ ,  $z_{EB}=0$ ,  $a=6378137$ ,  
 $f=0.0033528106647474805$ ,  $R_{Ee}=None$ )

Return the delta vector from position A to B.

##### Parameters

**n\_EA\_E, n\_EB\_E: 3 x n array**

n-vector(s) [no unit] of position A and B, decomposed in E.

**z\_EA, z\_EB: 1 x n array** Depth(s) [m] of system A and B, relative to the ellipsoid.  
 $(z_{EA} = \text{-height}, z_{EB} = \text{-height})$

**a: real scalar, default WGS-84 ellipsoid.** Semi-major axis of the Earth ellipsoid given in [m].

**f: real scalar, default WGS-84 ellipsoid.** Flattening [no unit] of the Earth ellipsoid. If  $f==0$  then spherical Earth with radius a is used in stead of WGS-84.

**R\_Ee** [3 x 3 array] rotation matrix defining the axes of the coordinate frame E.

#### Returns

**p\_AB\_E: 3 x n array** Cartesian position vector(s) from A to B, decomposed in E.

#### 1.13.1.1.8 Notes

The n-vectors for positions A ( $n_{EA\_E}$ ) and B ( $n_{EB\_E}$ ) are given. The output is the delta vector from A to B ( $p_{AB\_E}$ ). The calculation is exact, taking the ellipsity of the Earth into account. It is also non-singular as both n-vector and p-vector are non-singular (except for the center of the Earth). The default ellipsoid model used is WGS-84, but other ellipsoids/spheres might be specified.

#### 1.13.1.1.9 nvector.\_core.n\_EA\_E\_and\_p\_AB\_E2n\_EB\_E

**n\_EA\_E\_and\_p\_AB\_E2n\_EB\_E** ( $n_{EA\_E}$ ,  $p_{AB\_E}$ ,  $z_{EA}=0$ ,  $a=6378137$ ,  
 $f=0.0033528106647474805$ ,  $R_Ee=None$ )

Return position B from position A and delta.

#### Parameters

**n\_EA\_E: 3 x n array** n-vector(s) [no unit] of position A, decomposed in E.

**p\_AB\_E: 3 x n array** Cartesian position vector(s) from A to B, decomposed in E.

**z\_EA: 1 x n array** Depth(s) [m] of system A, relative to the ellipsoid. ( $z_{EA} = -\text{height}$ )

**a: real scalar, default WGS-84 ellipsoid.** Semi-major axis of the Earth ellipsoid given in [m].

**f: real scalar, default WGS-84 ellipsoid.** Flattening [no unit] of the Earth ellipsoid. If  $f==0$  then spherical Earth with radius a is used in stead of WGS-84.

**R\_Ee** [3 x 3 array] rotation matrix defining the axes of the coordinate frame E.

#### Returns

**n\_EB\_E: 3 x n array** n-vector(s) [no unit] of position B, decomposed in E.

**z\_EB: 1 x n array** Depth(s) [m] of system B, relative to the ellipsoid. ( $z_{EB} = -\text{height}$ )

#### See also:

[n\\_EA\\_E\\_and\\_n\\_EB\\_E2p\\_AB\\_E](#), [p\\_EB\\_E2n\\_EB\\_E](#), [n\\_EB\\_E2p\\_EB\\_E](#)

#### 1.13.1.1.10 Notes

The n-vector for position A ( $n_{EA\_E}$ ) and the position-vector from position A to position B ( $p_{AB\_E}$ ) are given. The output is the n-vector of position B ( $n_{EB\_E}$ ) and depth of B ( $z_{EB}$ ). The calculation is exact, taking the ellipsity of the Earth into account. It is also non-singular as both n-vector and p-vector are non-singular (except for the center of the Earth). The default ellipsoid model used is WGS-84, but other ellipsoids/spheres might be specified.

### 1.13.1.1.11 nvector.core.n\_EA\_E\_and\_n\_EB\_E2azimuth

**n\_EA\_E\_and\_n\_EB\_E2azimuth** (*n\_EA\_E*, *n\_EB\_E*, *a*=6378137, *f*=0.0033528106647474805,  
*R\_Ee=None*)

Return azimuth from A to B, relative to North:

#### Parameters

**n\_EA\_E, n\_EB\_E: 3 x n array** n-vector(s) [no unit] of position A and B, respectively, decomposed in E.

**a: real scalar, default WGS-84 ellipsoid.** Semi-major axis of the Earth ellipsoid given in [m].

**f: real scalar, default WGS-84 ellipsoid.** Flattening [no unit] of the Earth ellipsoid. If *f*=0 then spherical Earth with radius *a* is used in stead of WGS-84.

**R\_Ee** [3 x 3 array] rotation matrix defining the axes of the coordinate frame E.

#### Returns

**azimuth: n array** Angle [rad] the line makes with a meridian, taken clockwise from north.

### 1.13.1.1.12 nvector.core.n\_EA\_E\_distance\_and\_azimuth2n\_EB\_E

**n\_EA\_E\_distance\_and\_azimuth2n\_EB\_E** (*n\_EA\_E*, *distance\_rad*, *azimuth*, *R\_Ee=None*)

Return position B from azimuth and distance from position A

#### Parameters

**n\_EA\_E: 3 x n array**

n-vector(s) [no unit] of position A decomposed in E.

**distance\_rad: n, array** great circle distance [rad] from position A to B

**azimuth: n array** Angle [rad] the line makes with a meridian, taken clockwise from north.

#### Returns

**n\_EB\_E: 3 x n array** n-vector(s) [no unit] of position B decomposed in E.

### 1.13.1.1.13 nvector.core.great\_circle\_distance

**great\_circle\_distance** (*n\_EA\_E*, *n\_EB\_E*, *radius*=6371009.0)

Return great circle distance between positions A and B

#### Parameters

**n\_EA\_E, n\_EB\_E: 3 x n array**

n-vector(s) [no unit] of position A and B, decomposed in E.

**radius: real scalar** radius of sphere.

Formulae is given by equation (16) in Gade (2010) and is well conditioned for all angles.

### 1.13.1.1.14 nvector.\_core.euclidean\_distance

**euclidean\_distance** (*n\_EA\_E*, *n\_EB\_E*, *radius*=6371009.0)

Return Euclidean distance between positions A and B

#### Parameters

**n\_EA\_E**, **n\_EB\_E**: 3 x n array

n-vector(s) [no unit] of position A and B, decomposed in E.

**radius: real scalar** radius of sphere.

### 1.13.1.1.15 nvector.\_core.cross\_track\_distance

**cross\_track\_distance** (*path*, *n\_EB\_E*, *method*='greatcircle', *radius*=6371009.0)

Return cross track distance between path A and position B.

#### Parameters

**path: tuple of 2 n-vectors**

2 n-vectors of positions defining path A, decomposed in E.

**n\_EB\_E: 3 x m array** n-vector(s) of position B to measure the cross track distance to.

**method: string** defining distance calculated. Options are: 'greatcircle' or 'euclidean'

**radius: real scalar** radius of sphere. (default 6371009.0)

#### Returns

**distance** [array of length max(n, m)] cross track distance(s)

### 1.13.1.1.16 nvector.\_core.closest\_point\_on\_great\_circle

**closest\_point\_on\_great\_circle** (*path*, *n\_EB\_E*)

Return closest point C on great circle path A to position B.

#### Parameters

**path: tuple of 2 n-vectors of 3 x n arrays**

2 n-vectors of positions defining path A, decomposed in E.

**n\_EB\_E: 3 x m array** n-vector(s) of position B to find the closest point to.

#### Returns

**n\_EC\_E: 3 x max(m, n) array** n-vector(s) of closest position C on great circle path A

### 1.13.1.1.17 nvector.\_core.intersect

**intersect** (*path\_a*, *path\_b*)

Return the intersection(s) between the great circles of the two paths

#### Parameters

**path\_a, path\_b: tuple of 2 n-vectors** defining path A and path B, respectively. Path A and B has shape 2 x 3 x n and 2 x 3 x m, respectively.

#### Returns

---

**n\_EC\_E** [array of shape 3 x max(n, m)] n-vector(s) [no unit] of position C decomposed in E. point(s) of intersection between paths.

### 1.13.1.1.18 nvector\_core.mean\_horizontal\_position

**mean\_horizontal\_position** (*n\_EB\_E*)

Return the n-vector of the horizontal mean position.

#### Parameters

**n\_EB\_E: 3 x n array** n-vectors [no unit] of positions Bi, decomposed in E.

#### Returns

**p\_EM\_E: 3 x 1 array** n-vector [no unit] of the mean positions of all Bi, decomposed in E.

### 1.13.1.1.19 nvector\_core.on\_great\_circle

**on\_great\_circle** (*path, n\_EB\_E, radius=6371009.0, rtol=1e-06, atol=1e-08*)

True if position B is on great circle through path A.

#### Parameters

**path: tuple of 2 n-vectors**

2 n-vectors of positions defining path A, decomposed in E.

**n\_EB\_E: 3 x m array** n-vector(s) of position B to check to.

**radius: real scalar** radius of sphere. (default 6371009.0)

**rtol, atol: real scalars** defining relative and absolute tolerance

#### Returns

**on** [bool array of length max(n, m)] True if position B is on great circle through path A.

### 1.13.1.1.20 nvector\_core.on\_great\_circle\_path

**on\_great\_circle\_path** (*path, n\_EB\_E, radius=6371009.0, rtol=1e-06, atol=1e-08*)

True if position B is on great circle and between endpoints of path A.

#### Parameters

**path: tuple of 2 n-vectors**

2 n-vectors of positions defining path A, decomposed in E.

**n\_EB\_E: 3 x m array** n-vector(s) of position B to measure the cross track distance to.

**radius: real scalar** radius of sphere. (default 6371009.0)

**rtol, atol: real scalars** defining relative and absolute tolerance

#### Returns

**on** [bool array of length max(n, m)] True if position B is on great circle and between end-points of path A.

## 1.13.1.2 Rotation matrices and angles

<code>E_rotation([axes])</code>	Return rotation matrix R_Ee defining the axes of the coordinate frame E.
<code>n_E2R_EN(n_E[, R_Ee])</code>	Returns the rotation matrix R_EN from n-vector.
<code>n_E_and_wa2R_EL(n_E, wander_azimuth[, R_Ee])</code>	Returns rotation matrix R_EL from n-vector and wander azimuth angle.
<code>R_EL2n_E(R_EL)</code>	Returns n-vector from the rotation matrix R_EL.
<code>R_EN2n_E(R_EN)</code>	Returns n-vector from the rotation matrix R_EN.
<code>R2xyz(R_AB)</code>	Returns the angles about new axes in the xyz-order from a rotation matrix.
<code>R2zyx(R_AB)</code>	Returns the angles about new axes in the zxy-order from a rotation matrix.
<code>xyz2R(x, y, z)</code>	Returns rotation matrix from 3 angles about new axes in the xyz-order.
<code>zyx2R(z, y, x)</code>	Returns rotation matrix from 3 angles about new axes in the zyx-order.

### 1.13.1.2.1 nvector\_core.E\_rotation

`E_rotation(axes='e')`

Return rotation matrix R\_Ee defining the axes of the coordinate frame E.

#### Parameters

`axes` ['e' or 'E'] defines orientation of the axes of the coordinate frame E. Options are: 'e': z-axis points to the North Pole along the Earth's rotation axis,

x-axis points towards the point where latitude = longitude = 0. This choice is very common in many fields.

**'E': x-axis points to the North Pole along the Earth's rotation axis,** y-axis points towards longitude +90deg (east) and latitude = 0. (the yz-plane coincides with the equatorial plane). This choice of axis ensures that at zero latitude and longitude, frame N (North-East-Down) has the same orientation as frame E. If roll/pitch/yaw are zero, also frame B (forward-starboard-down) has this orientation. In this manner, the axes of frame E is chosen to correspond with the axes of frame N and B. The functions in this library originally used this option.

#### Returns

`R_Ee` [3 x 3 array] rotation matrix defining the axes of the coordinate frame E as described in Table 2 in Gade (2010)

**R\_Ee controls the axes of the coordinate frame E (Earth-Centred, Earth-Fixed, ECEF) used by the other functions in this library**

### 1.13.1.2.2 Examples

```
>>> import numpy as np
>>> import nvector as nv
>>> np.allclose(nv.E_rotation(axes='e'), [[ 0.,  0.,  1.],
...                                         [ 0.,  1.,  0.],
...                                         [-1.,  0.,  0.]])
True
>>> np.allclose(nv.E_rotation(axes='E'), [[ 1.,  0.,  0.],
...                                         [ 0.,  1.,  0.],
...                                         [ 0.,  0.,  1.]])
True
```

### 1.13.1.2.3 nvector\_core.n\_E2R\_EN

**n\_E2R\_EN** (*n\_E*, *R\_Ee=None*)

Returns the rotation matrix R\_EN from n-vector.

#### Parameters

**n\_E: 3 x n array** n-vector [no unit] decomposed in E

**R\_Ee** [3 x 3 array] rotation matrix defining the axes of the coordinate frame E.

#### Returns

**R\_EN: 3 x 3 x n array** The resulting rotation matrix [no unit] (direction cosine matrix).

See also:

[R\\_EN2n\\_E](#), [n\\_E\\_and\\_wa2R\\_EL](#), [R\\_EL2n\\_E](#)

### 1.13.1.2.4 nvector\_core.n\_E\_and\_wa2R\_EL

**n\_E\_and\_wa2R\_EL** (*n\_E*, *wander\_azimuth*, *R\_Ee=None*)

Returns rotation matrix R\_EL from n-vector and wander azimuth angle.

**R\_EL = n\_E\_and\_wa2R\_EL(n\_E, wander\_azimuth)** Calculates the rotation matrix (direction cosine matrix) R\_EL using n-vector (*n\_E*) and the wander azimuth angle. When *wander\_azimuth=0*, we have that N=L (See Table 2 in Gade (2010) for details)

#### Parameters

**n\_E: 3 x n array** n-vector [no unit] decomposed in E

**wander\_azimuth: real scalar or array of length n** Angle [rad] between L's x-axis and north, positive about L's z-axis.

**R\_Ee** [3 x 3 array] rotation matrix defining the axes of the coordinate frame E.

#### Returns

**R\_EL: 3 x 3 x n array** The resulting rotation matrix. [no unit]

See also:

[R\\_EL2n\\_E](#), [R\\_EN2n\\_E](#), [n\\_E2R\\_EN](#)

### 1.13.1.2.5 nvector\_core.R\_EL2n\_E

**R\_EL2n\_E** (*R\_EL*)

Returns n-vector from the rotation matrix R\_EL.

#### Parameters

**R\_EL: 3 x 3 x n array** Rotation matrix (direction cosine matrix) [no unit]

#### Returns

**n\_E: 3 x n array** n-vector(s) [no unit] decomposed in E.

See also:

[R\\_EN2n\\_E](#), [n\\_E\\_and\\_wa2R\\_EL](#), [n\\_E2R\\_EN](#)

### 1.13.1.2.6 nvector\_core.R\_EN2n\_E

**R\_EN2n\_E** ( $R_{EN}$ )

Returns n-vector from the rotation matrix R\_EN.

#### Parameters

**R\_EN: 3 x 3 x n array** Rotation matrix (direction cosine matrix) [no unit]

#### Returns

**n\_E: 3 x n array** n-vector [no unit] decomposed in E.

See also:

[n\\_E2R\\_EN](#), [R\\_EL2n\\_E](#), [n\\_E\\_and\\_wa2R\\_EL](#)

### 1.13.1.2.7 nvector\_core.R2xyz

**R2xyz** ( $R_{AB}$ )

Returns the angles about new axes in the xyz-order from a rotation matrix.

#### Parameters

**R\_AB: 3 x 3 x n array** rotation matrix [no unit] (direction cosine matrix) such that the relation between a vector v decomposed in A and B is given by:  $v_A = \text{mdot}(R_{AB}, v_B)$

#### Returns

**x, y, z: real scalars or array of length n.** Angles [rad] of rotation about new axes.

See also:

[xyz2R](#), [R2zyx](#), [xyz2R](#)

### 1.13.1.2.8 Notes

The x, y, z angles are called Euler angles or Tait-Bryan angles and are defined by the following procedure of successive rotations: Given two arbitrary coordinate frames A and B. Consider a temporary frame T that initially coincides with A. In order to make T align with B, we first rotate T an angle x about its x-axis (common axis for both A and T). Secondly, T is rotated an angle y about the NEW y-axis of T. Finally, T is rotated an angle z about its NEWEST z-axis. The final orientation of T now coincides with the orientation of B.

The signs of the angles are given by the directions of the axes and the right hand rule.

### 1.13.1.2.9 nvector\_core.R2zyx

**R2zyx** ( $R_{AB}$ )

Returns the angles about new axes in the zxy-order from a rotation matrix.

#### Parameters

**R\_AB: 3x3 array** rotation matrix [no unit] (direction cosine matrix) such that the relation between a vector v decomposed in A and B is given by:  $v_A = \text{np.dot}(R_{AB}, v_B)$

#### Returns

**z, y, x: real scalars** Angles [rad] of rotation about new axes.

See also:

[zyx2R](#), [xyz2R](#), [R2xyz](#)

### 1.13.1.2.10 Notes

The z, x, y angles are called Euler angles or Tait-Bryan angles and are defined by the following procedure of successive rotations: Given two arbitrary coordinate frames A and B. Consider a temporary frame T that initially coincides with A. In order to make T align with B, we first rotate T an angle z about its z-axis (common axis for both A and T). Secondly, T is rotated an angle y about the NEW y-axis of T. Finally, T is rotated an angle x about its NEWEST x-axis. The final orientation of T now coincides with the orientation of B.

The signs of the angles are given by the directions of the axes and the right hand rule.

Note that if A is a north-east-down frame and B is a body frame, we have that z=yaw, y=pitch and x=roll.

### 1.13.1.2.11 nvector.\_core.xyz2R

**xyz2R** (*x, y, z*)

Returns rotation matrix from 3 angles about new axes in the xyz-order.

#### Parameters

**x,y,z: real scalars or array of lengths n** Angles [rad] of rotation about new axes.

#### Returns

**R\_AB: 3 x 3 x n array** rotation matrix [no unit] (direction cosine matrix) such that the relation between a vector v decomposed in A and B is given by:  $v_A = \text{mdot}(R_{AB}, v_B)$

See also:

[R2xyz](#), [zyx2R](#), [R2zyx](#)

### 1.13.1.2.12 Notes

The rotation matrix R\_AB is created based on 3 angles x,y,z about new axes (intrinsic) in the order x-y-z. The angles are called Euler angles or Tait-Bryan angles and are defined by the following procedure of successive rotations: Given two arbitrary coordinate frames A and B. Consider a temporary frame T that initially coincides with A. In order to make T align with B, we first rotate T an angle x about its x-axis (common axis for both A and T). Secondly, T is rotated an angle y about the NEW y-axis of T. Finally, T is rotated an angle z about its NEWEST z-axis. The final orientation of T now coincides with the orientation of B.

The signs of the angles are given by the directions of the axes and the right hand rule.

### 1.13.1.2.13 nvector.\_core.zyx2R

**zyx2R** (*z, y, x*)

Returns rotation matrix from 3 angles about new axes in the zyx-order.

#### Parameters

**z, y, x: real scalars or arrays of lengths n** Angles [rad] of rotation about new axes.

#### Returns

**R\_AB: 3 x 3 x n array** rotation matrix [no unit] (direction cosine matrix) such that the relation between a vector v decomposed in A and B is given by:  $v_A = \text{mdot}(R_{AB}, v_B)$

See also:

[R2zyx](#), [xyz2R](#), [R2xyz](#)

### 1.13.1.2.14 Notes

The rotation matrix  $R_{AB}$  is created based on 3 angles z,y,x about new axes (intrinsic) in the order z-y-x. The angles are called Euler angles or Tait-Bryan angles and are defined by the following procedure of successive rotations: Given two arbitrary coordinate frames A and B. Consider a temporary frame T that initially coincides with A. In order to make T align with B, we first rotate T an angle z about its z-axis (common axis for both A and T). Secondly, T is rotated an angle y about the NEW y-axis of T. Finally, T is rotated an angle x about its NEWEST x-axis. The final orientation of T now coincides with the orientation of B.

The signs of the angles are given by the directions of the axes and the right hand rule.

Note that if A is a north-east-down frame and B is a body frame, we have that z=yaw, y=pitch and x=roll.

### 1.13.1.3 Misc functions

<code>nthroot(x, n)</code>	Return the n'th root of x to machine precision
<code>deg(rad_angle)</code>	Converts angle in radians to degrees.
<code>rad(deg_angle)</code>	Converts angle in degrees to radians.
<code>select_ellipsoid(name)</code>	Return semi-major axis (a), flattening (f) and name of ellipsoid
<code>unit(vector[, norm_zero_vector])</code>	Convert input vector to a vector of unit length.

#### 1.13.1.3.1 nvector\_core.nthroot

`nthroot (x, n)`

Return the n'th root of x to machine precision

Parameters x, n

#### 1.13.1.3.2 Examples

```
>>> import numpy as np
>>> import nvector as nv
>>> np.allclose(nv.nthroot(27.0, 3), 3.0)
True
```

#### 1.13.1.3.3 nvector\_core.deg

`deg (rad_angle)`

Converts angle in radians to degrees.

**Parameters**

`rad_angle:` angle in radians

**Returns**

`deg_angle:` angle in degrees

**See also:**

`rad`

### 1.13.1.3.4 nvector\_core.rad

**rad**(*deg\_angle*)

Converts angle in degrees to radians.

#### Parameters

**deg\_angle:** angle in degrees

#### Returns

**rad\_angle:** angle in radians

See also:

*deg*

### 1.13.1.3.5 nvector\_core.select\_ellipsoid

**select\_ellipsoid**(*name*)

Return semi-major axis (a), flattening (f) and name of ellipsoid

#### Parameters

**name** [string] name of ellipsoid. Valid options are: 1) Airy 1858 2) Airy Modified 3) Australian National 4) Bessel 1841 5) Clarke 1880 6) Everest 1830 7) Everest Modified 8) Fisher 1960 9) Fisher 1968 10) Hough 1956 11) International (Hayford) 12) Krassovsky 1938 13) NWL-9D (WGS 66) 14) South American 1969 15) Soviet Geod. System 1985 16) WGS 72 17) Clarke 1866 (NAD27) 18) GRS80 / WGS84 (NAD83)

### 1.13.1.3.6 Examples

```
>>> import nvector as nv
>>> nv.select_ellipsoid(name='wgs84')
(6378137.0, 0.0033528106647474805, 'GRS80 / WGS84 (NAD83)')
>>> nv.select_ellipsoid(name='GRS80')
(6378137.0, 0.0033528106647474805, 'GRS80 / WGS84 (NAD83)')
>>> nv.select_ellipsoid(name='NAD83')
(6378137.0, 0.0033528106647474805, 'GRS80 / WGS84 (NAD83)')
>>> nv.select_ellipsoid(name=18)
(6378137.0, 0.0033528106647474805, 'GRS80 / WGS84 (NAD83)')
```

### 1.13.1.3.7 nvector\_core.unit

**unit**(*vector*, *norm\_zero\_vector=1*)

Convert input vector to a vector of unit length.

#### Parameters

**vector** [3 x m array] m column vectors

#### Returns

**unitvector** [3 x m array] normalized unitvector(s) along axis==0.

### 1.13.1.3.8 Notes

The column vector(s) that have zero length will be returned as unit vector(s) pointing in the x-direction, i.e, [[1], [0], [0]]

### 1.13.1.3.9 Examples

```
>>> import numpy as np
>>> import nvector as nv
>>> np.allclose(nv.unit([[1, 0], [1, 0], [1, 0]]), [[ 0.57735027,  1],
...                                         [ 0.57735027,  0],
...                                         [ 0.57735027,  0]])
True
```

### 1.13.1.4 OO interface to Geodesic functions

<code>FrameE([a, f, name, axes])</code>	Earth-fixed frame
<code>FrameN(position)</code>	North-East-Down frame
<code>FrameL(position[, wander_azimuth])</code>	Local level, Wander azimuth frame
<code>FrameB(position[, yaw, pitch, roll, degrees])</code>	Body frame
<code>ECEFvector(pvector[, frame])</code>	Geographical position given as Cartesian position vector in frame E
<code>GeoPoint(latitude, longitude[, z, frame, ...])</code>	Geographical position given as latitude, longitude, depth in frame E
<code>Nvector(normal[, z, frame])</code>	Geographical position given as n-vector and depth in frame E
<code>GeoPath(positionA, positionB)</code>	Geographical path between two positions in Frame E
<code>Pvector(pvector, frame)</code>	
<code>diff_positions(*args, **kwds)</code>	<i>diff_positions</i> is deprecated!

#### 1.13.1.4.1 nvector.objects.FrameE

`class FrameE(a=None, f=None, name='WGS84', axes='e')`

Earth-fixed frame

##### Parameters

**a:** real scalar, default WGS-84 ellipsoid.

Semi-major axis of the Earth ellipsoid given in [m].

**f:** real scalar, default WGS-84 ellipsoid.

Flattening [no unit] of the Earth ellipsoid. If  $f==0$  then spherical Earth with radius  $a$  is used in stead of WGS-84.

**name:** string defining the default ellipsoid.

**axes:** ‘e’ or ‘E’ defines axes orientation of E frame. Default is axes=’e’ which means that the orientation of the axis is such that: z-axis -> North Pole, x-axis -> Latitude=Longitude=0.

##### See also:

`FrameN`, `FrameL`, `FrameB`

#### 1.13.1.4.2 Notes

The frame is Earth-fixed (rotates and moves with the Earth) where the origin coincides with Earth’s centre (geometrical centre of ellipsoid model).

`__init__(a=None, f=None, name='WGS84', axes='e')`  
`x.__init__(...) initializes x; see help(type(x)) for signature`

#### 1.13.1.4.3 Methods

ECEFvector(*args, **kwds)	Geographical position given as Cartesian position vector in frame E
GeoPoint(*args, **kwds)	Geographical position given as latitude, longitude, depth in frame E
Nvector(*args, **kwds)	Geographical position given as n-vector and depth in frame E
<code>__init__</code> ([a, f, name, axes])	x. <code>__init__</code> (...) initializes x; see help(type(x)) for signature
direct(lat_a, lon_a, azimuth, distance[, z, ...])	Return position B computed from position A, distance and azimuth.
inverse(lat_a, lon_a, lat_b, lon_b[, z, ...])	Return ellipsoidal distance between positions as well as the direction.

#### 1.13.1.4.4 nvector.objects.FrameN

**class FrameN**(position)

North-East-Down frame

##### Parameters

**position:** ECEFvector, GeoPoint or Nvector object position of the vehicle (B) which also defines the origin of the local frame N. The origin is directly beneath or above the vehicle (B), at Earth's surface (surface of ellipsoid model).

#### 1.13.1.4.5 Notes

The Cartesian frame is local and oriented North-East-Down, i.e., the x-axis points towards north, the y-axis points towards east (both are horizontal), and the z-axis is pointing down.

When moving relative to the Earth, the frame rotates about its z-axis to allow the x-axis to always point towards north. When getting close to the poles this rotation rate will increase, being infinite at the poles. The poles are thus singularities and the direction of the x- and y-axes are not defined here. Hence, this coordinate frame is NOT SUITABLE for general calculations.

`__init__`(position)

x.`__init__`(...) initializes x; see help(type(x)) for signature

#### 1.13.1.4.6 Methods

Pvector(pvector)	
<code>__init__</code> (position)	x. <code>__init__</code> (...) initializes x; see help(type(x)) for signature

#### 1.13.1.4.7 Attributes

---

R\_EN

---

### 1.13.1.4.8 nvector.objects.FrameL

**class FrameL**(*position*, *wander\_azimuth*=0)

Local level, Wander azimuth frame

#### Parameters

**position:** ECEFvector, GeoPoint or Nvector object position of the vehicle (B) which also defines the origin of the local frame L. The origin is directly beneath or above the vehicle (B), at Earth's surface (surface of ellipsoid model).

**wander\_azimuth:** real scalar Angle [rad] between the x-axis of L and the north direction.

#### See also:

*FrameE*, *FrameN*, *FrameB*

### 1.13.1.4.9 Notes

The Cartesian frame is local and oriented Wander-azimuth-Down. This means that the z-axis is pointing down. Initially, the x-axis points towards north, and the y-axis points towards east, but as the vehicle moves they are not rotating about the z-axis (their angular velocity relative to the Earth has zero component along the z-axis).

(Note: Any initial horizontal direction of the x- and y-axes is valid for L, but if the initial position is outside the poles, north and east are usually chosen for convenience.)

The L-frame is equal to the N-frame except for the rotation about the z-axis, which is always zero for this frame (relative to E). Hence, at a given time, the only difference between the frames is an angle between the x-axis of L and the north direction; this angle is called the wander azimuth angle. The L-frame is well suited for general calculations, as it is non-singular.

**\_\_init\_\_**(*position*, *wander\_azimuth*=0)

x.\_\_init\_\_(...) initializes x; see help(type(x)) for signature

### 1.13.1.4.10 Methods

---

Pvector(pvector)

**\_\_init\_\_**(*position*[, *wander\_azimuth*])

x.\_\_init\_\_(...) initializes x; see help(type(x)) for signature

---

### 1.13.1.4.11 Attributes

---

R\_EN

---

### 1.13.1.4.12 nvector.objects.FrameB

**class FrameB**(*position*, *yaw*=0, *pitch*=0, *roll*=0, *degrees*=False)

Body frame

#### Parameters

**position:** ECEFvector, GeoPoint or Nvector object

position of the vehicle's reference point which also coincides with the origin of the frame B.

**yaw, pitch, roll: real scalars** defining the orientation of frame B in [deg] or [rad].

**degrees** [bool] if True yaw, pitch, roll are given in degrees otherwise in radians

#### 1.13.1.4.13 Notes

The frame is fixed to the vehicle where the x-axis points forward, the y-axis to the right (starboard) and the z-axis in the vehicle's down direction.

**\_\_init\_\_(position, yaw=0, pitch=0, roll=0, degrees=False)**  
x.\_\_init\_\_(...) initializes x; see help(type(x)) for signature

#### 1.13.1.4.14 Methods

---

Pvector(pvector)	
<b>__init__(position[, yaw, pitch, roll, degrees])</b>	x.__init__(...) initializes x; see help(type(x)) for signature

---

#### 1.13.1.4.15 Attributes

---

R_EN
------

---

#### 1.13.1.4.16 nvector.objects.ECEFvector

**class ECEFvector(pvector, frame=None)**

Geographical position given as Cartesian position vector in frame E

##### Parameters

**pvector: 3 x n array**

Cartesian position vector(s) [m] from E to B, decomposed in E.

**frame: FrameE object** reference ellipsoid. The default ellipsoid model used is WGS84, but other ellipsoids/spheres might be specified.

#### 1.13.1.4.17 Notes

The position of B (typically body) relative to E (typically Earth) is given into this function as p-vector, p\_EB\_E relative to the center of the frame.

**\_\_init\_\_(pvector, frame=None)**  
x.\_\_init\_\_(...) initializes x; see help(type(x)) for signature

#### 1.13.1.4.18 Methods

---

<b>__init__(pvector[, frame])</b>	x.__init__(...) initializes x; see help(type(x)) for signature
<b>change_frame(frame)</b>	Converts to Cartesian position vector in another frame
<b>delta_to(other)</b>	Return cartesian delta vector from positions A to B decomposed in N.

---

Continued on next page

Table 12 – continued from previous page

<code>to_ecef_vector()</code>	
<code>to_geo_point()</code>	Converts ECEF-vector to geo-point.
<code>to_nvector()</code>	Converts ECEF-vector to n-vector.

### 1.13.1.4.19 Attributes

<code>azimuth</code>	
<code>azimuth_deg</code>	
<code>elevation</code>	
<code>elevation_deg</code>	
<code>length</code>	

### 1.13.1.4.20 nvector.objects.GeoPoint

`class GeoPoint(latitude, longitude, z=0, frame=None, degrees=False)`

Geographical position given as latitude, longitude, depth in frame E

#### Parameters

- latitude, longitude:** real scalars or vectors of length n. Geodetic latitude and longitude given in [rad or deg]
- z:** real scalar or vector of length n. Depth(s) [m] relative to the ellipsoid (depth = -height)
- frame:** FrameE object reference ellipsoid. The default ellipsoid model used is WGS84, but other ellipsoids/spheres might be specified.
- degrees:** bool True if input are given in degrees otherwise radians are assumed.

### 1.13.1.4.21 Examples

Solve geodesic problems.

The following illustrates its use

```
>>> import nvector as nv
>>> wgs84 = nv.FrameE(name='WGS84')
```

The geodesic inverse problem

```
>>> positionA = wgs84.GeoPoint(-41.32, 174.81, degrees=True)
>>> positionB = wgs84.GeoPoint(40.96, -5.50, degrees=True)
>>> s12, az1, az2 = positionA.distance_and_azimuth(positionB, degrees=True)
>>> 's12 = {:.5.2f}, az1 = {:.5.2f}, az2 = {:.5.2f}'.format(s12, az1, az2)
's12 = 19959679.27, az1 = 161.07, az2 = 18.83'
```

The geodesic direct problem

```
>>> positionA = wgs84.GeoPoint(40.6, -73.8, degrees=True)
>>> az1, distance = 45, 10000e3
>>> positionB, az2 = positionA.displace(distance, az1, degrees=True)
>>> lat2, lon2 = positionB.latitude_deg, positionB.longitude_deg
>>> msg = 'lat2 = {:.5.2f}, lon2 = {:.5.2f}, az2 = {:.5.2f}'
>>> msg.format(lat2, lon2, az2)
'lat2 = 32.64, lon2 = 49.01, az2 = 140.37'
```

`__init__(latitude, longitude, z=0, frame=None, degrees=False)`

x.`__init__(...)` initializes x; see help(type(x)) for signature

#### 1.13.1.4.22 Methods

<code>__init__(latitude, longitude[, z, frame, ...])</code>	x. <code>__init__(...)</code> initializes x; see help(type(x)) for signature
<code>delta_to(other)</code>	Return cartesian delta vector from positions A to B decomposed in N.
<code>displace(distance, azimuth[, long_unroll, ...])</code>	Return position B computed from current position, distance and azimuth.
<code>distance_and_azimuth(point[, long_unroll, ...])</code>	Return ellipsoidal distance between positions as well as the direction.
<code>to_ecef_vector()</code>	Converts latitude and longitude to ECEF-vector.
<code>to_geo_point()</code>	Return geo-point
<code>to_nvector()</code>	Converts latitude and longitude to n-vector.

#### 1.13.1.4.23 Attributes

<code>latitude_deg</code>
<code>latlon</code>
<code>latlon_deg</code>
<code>longitude_deg</code>

#### 1.13.1.4.24 nvector.objects.Nvector

**class Nvector(normal, z=0, frame=None)**

Geographical position given as n-vector and depth in frame E

##### Parameters

**normal:** 3 x n array n-vector(s) [no unit] decomposed in E.

**z:** real scalar or vector of length n. Depth(s) [m] relative to the ellipsoid (depth = -height)

**frame:** FrameE object reference ellipsoid. The default ellipsoid model used is WGS84, but other ellipsoids/spheres might be specified.

##### See also:

`GeoPoint`, `ECEFvector`, `Pvector`

#### 1.13.1.4.25 Notes

The position of B (typically body) relative to E (typically Earth) is given into this function as n-vector, n\_EB\_E and a depth, z relative to the ellipsoid.

`__init__(normal, z=0, frame=None)`

x.`__init__(...)` initializes x; see help(type(x)) for signature

#### 1.13.1.4.26 Methods

<code>__init__(normal[, z, frame])</code>	x. <code>__init__(...)</code> initializes x; see help(type(x)) for signature
<code>delta_to(other)</code>	Return cartesian delta vector from positions A to B decomposed in N.
<code>mean()</code>	Return mean position of the n-vectors.

Continued on next page

Table 16 – continued from previous page

<code>mean_horizontal_position(**kwds)</code>	<i>mean_horizontal_position</i> is deprecated!
<code>to_ecef_vector()</code>	Converts n-vector to Cartesian position vector (“ECEF-vector”)
<code>to_geo_point()</code>	Converts n-vector to geo-point.
<code>to_nvector()</code>	
<code>unit()</code>	Normalizes self to unit vector(s)

#### 1.13.1.4.27 nvector.objects.GeoPath

**class GeoPath(*positionA, positionB*)**

Geographical path between two positions in Frame E

##### Parameters

**positionA, positionB: Nvector, GeoPoint or ECEFvector objects** The path is defined by the line between position A and B, decomposed in E.

`__init__(positionA, positionB)`

*x.\_\_init\_\_(...)* initializes *x*; see `help(type(x))` for signature

#### 1.13.1.4.28 Methods

<code>__init__(<i>positionA, positionB</i>)</code>	<i>x.__init__(...)</i> initializes <i>x</i> ; see <code>help(type(x))</code> for signature
<code>closest_point_on_great_circle(<i>point</i>)</code>	
<code>closest_point_on_path(<i>point</i>)</code>	Returns closest point on great circle path segment to the point.
<code>cross_track_distance(<i>point</i>[, <i>method</i>, <i>radius</i>])</code>	Return cross track distance from path to point.
<code>ecef_vectors()</code>	Return <i>positionA</i> and <i>positionB</i> as ECEF-vectors
<code>geo_points()</code>	Return <i>positionA</i> and <i>positionB</i> as geo-points
<code>interpolate(<i>ti</i>)</code>	Return the interpolated point along the path
<code>intersect(<i>path</i>)</code>	Return the intersection(s) between the great circles of the two paths
<code>intersection(**kwds)</code>	<i>intersection</i> is deprecated!
<code>nvector_normals()</code>	
<code>nvectors()</code>	Return <i>positionA</i> and <i>positionB</i> as n-vectors
<code>on_great_circle(<i>point</i>[, <i>rtol</i>, <i>atol</i>])</code>	
<code>on_path(<i>point</i>[, <i>method</i>, <i>rtol</i>, <i>atol</i>])</code>	Return True if <i>point</i> is on the path between A and B
<code>track_distance([<i>method</i>, <i>radius</i>])</code>	Return the distance of the path.

#### 1.13.1.4.29 nvector.objects.Pvector

**class Pvector(*pvector, frame*)**

`__init__(pvector, frame)`

*x.\_\_init\_\_(...)* initializes *x*; see `help(type(x))` for signature

#### 1.13.1.4.30 Methods

<code>__init__(pvector, frame)</code>	x.__init__(...) initializes x; see help(type(x)) for signature
<code>delta_to(other)</code>	Return cartesian delta vector from positions A to B decomposed in N.
<code>to_ecef_vector()</code>	
<code>to_geo_point()</code>	
<code>to_nvector()</code>	

#### 1.13.1.4.31 Attributes

<code>azimuth</code>
<code>azimuth_deg</code>
<code>elevation</code>
<code>elevation_deg</code>
<code>length</code>

#### 1.13.1.4.32 nvector.objects.diff\_positions

`diff_positions(*args, **kwds)`

*diff\_positions* is deprecated!

Deprecated use `delta_E` instead.



# CHAPTER 2

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